

**ONE HOUR THIRTY MINUTES**

A list of constants is enclosed.

**UNIVERSITY OF MANCHESTER**

Vector Spaces for Quantum Mechanics

1st June 2011, 9.45 a.m. - 11.15 a.m.

Answer **ALL** parts of question 1 and **TWO** other questions

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Electronic calculators may be used, provided that they cannot store text.

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The numbers are given as a guide to the relative weights of the different parts of each question.

1. a) The kets  $|a\rangle$  and  $|b\rangle$  are represented in a certain orthonormal basis by  $\begin{pmatrix} -2 \\ 2i \end{pmatrix}$  and  $\begin{pmatrix} 2 + 3i \\ 2i \end{pmatrix}$  respectively. Find the numerical values of  $\langle a|b\rangle$  and  $\langle b|a\rangle$ , and the norm of  $|c\rangle = |a\rangle + |b\rangle$ .

[6 marks]

- b) For each of the following matrices, state whether it is Hermitian, unitary, both or neither:

$$(i) \begin{pmatrix} 1 & i \\ -i & 1 \end{pmatrix}; \quad (ii) \begin{pmatrix} i & 1 \\ 1 & -i \end{pmatrix}; \quad (iii) \frac{1}{\sqrt{2}} \begin{pmatrix} 1 & i \\ 1 & -i \end{pmatrix}; \quad (iv) \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}.$$

[6 marks]

- c) How many dimensions are needed for the vector space describing the combined spins of an electron and a proton (both spin- $\frac{1}{2}$ )? Write down an orthonormal basis for this space in terms of the single-particle basis vectors  $\{|\uparrow\rangle_e, |\downarrow\rangle_e\}$  and  $\{|\uparrow\rangle_p, |\downarrow\rangle_p\}$  for the electron and proton respectively. Write down the possible eigenvalues of the total-spin operator  $\hat{S}^2$  for the two-particle system. Without calculation, say which of your two-particle basis vectors are eigenvectors of  $\hat{S}^2$ .

[7 marks]

- d) Let  $\{|x\rangle\}$  be the set of eigenkets of the position operator on a function space describing one-dimensional functions of  $x$ , normalised according to the delta-function convention.

(i) Given that  $|f\rangle, |g\rangle$  in the function space represent the functions  $f(x), g(x)$ , write down  $\langle f|g\rangle$  in terms of  $f(x)$  and  $g(x)$ .

(ii) Write down the expansion for the identity operator  $\hat{I}$  in terms of the  $\{|x\rangle\}$ .

(iii) Using your answer to (ii), show that  $\langle f|\hat{I}|g\rangle$  is the same as your answer to (i).

[6 marks]

2. a) Given an operator  $\hat{A}$ , write down an expression for a matrix element  $A_{jk}$  in a particular basis, say  $\{|y_i\rangle\}$ . Hence, for a spin- $\frac{1}{2}$  system, find the matrix for the  $\hat{S}_y$  operator, in the basis of its eigenkets  $\{|+y\rangle, |-y\rangle\}$  (representing spin “up” and “down” along the  $y$  axis).

[7 marks]

- b) In terms of the eigenkets of  $\hat{S}_z$ , i.e.  $\{|\uparrow\rangle, |\downarrow\rangle\}$ , we can write the  $\hat{S}_y$  eigenkets as:

$$|+y\rangle = \frac{|\uparrow\rangle + i|\downarrow\rangle}{\sqrt{2}}, \quad |-y\rangle = \frac{|\uparrow\rangle - i|\downarrow\rangle}{\sqrt{2}}.$$

Construct the matrix of eigenvectors of  $\hat{S}_y$  in the  $S_z$  basis, and use it to transform  $|\uparrow\rangle, |\downarrow\rangle$ , and the  $\hat{S}_z$  operator from the  $S_z$  to the  $S_y$  representation. You may assume that in the  $S_z$  representation

$$\hat{S}_z \xrightarrow{S_z} \frac{\hbar}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}.$$

[12 marks]

- c) Evaluate the matrix of  $\hat{A} = [\hat{S}_y, \hat{S}_z]$  using your  $S_y$ -representation matrices. What observable is represented by  $\hat{A}/i\hbar$ ?

[6 marks]

3. a) The lowering operator  $\hat{a}$  for a quantum simple harmonic oscillator (SHO) can be written:

$$\hat{a} = \sqrt{\frac{m\omega}{2\hbar}} \left( \hat{x} + \frac{i}{m\omega} \hat{p} \right)$$

Show that the product  $\hat{a}^\dagger \hat{a}$  is equal to the number operator  $\hat{N}$ , whose eigenvalues are the energy level quantum numbers  $n$ :

$$E_n = \left( n + \frac{1}{2} \right) \hbar\omega.$$

[7 marks]

- b) Let  $\{|n\rangle\}_{n=0}^{\infty}$  be the usual energy eigenstates of the SHO. A so-called coherent state of a quantum SHO is given by

$$|\lambda\rangle = A \sum_{n=0}^{\infty} \frac{\lambda^n}{\sqrt{n!}} |n\rangle,$$

where  $A$  is a normalising constant and  $\lambda$  is any complex number. Show that  $|\lambda\rangle$  is an eigenstate of the lowering operator  $\hat{a}$ , with eigenvalue  $\lambda$ .

[8 marks]

- c) Evaluate  $\langle \lambda | \hat{a}^\dagger \hat{a} | \lambda \rangle$  and  $\langle \lambda | (\hat{a}^\dagger \hat{a})^2 | \lambda \rangle$ , and hence show that for state  $|\lambda\rangle$ , the expected value of the energy quantum number is equal to the square of its uncertainty:  $\langle n \rangle = (\Delta n)^2$ .

[10 marks]

You may assume that  $\hat{a}|n\rangle = \sqrt{n}|n-1\rangle$ , and that  $[\hat{a}, \hat{a}^\dagger] = 1$ .

4. A system has orbital angular momentum quantum number  $l = 1$ . In the  $L_z$  basis,  $\hat{L}_y$  is represented by

$$\hat{L}_y \xrightarrow{L_z} \frac{\hbar}{\sqrt{2}} \begin{pmatrix} 0 & -i & 0 \\ i & 0 & -i \\ 0 & i & 0 \end{pmatrix}.$$

a) The system is in a magnetic field aligned with the z-axis, and the Hamiltonian can be written  $\widehat{H} = -\mu\hat{L}_z$ . If the system is originally in the state with  $L_y = 0$ , i.e.

$$|\psi(t=0)\rangle = |L_y = 0\rangle \xrightarrow{L_z} \frac{1}{\sqrt{2}} \begin{pmatrix} 1 \\ 0 \\ 1 \end{pmatrix},$$

show that the state at time  $t$  is represented in the  $L_z$  basis by

$$|\psi(t)\rangle \xrightarrow{L_z} \frac{1}{\sqrt{2}} \begin{pmatrix} e^{i\mu t} \\ 0 \\ e^{-i\mu t} \end{pmatrix}$$

[6 marks]

b) Hence find  $\langle L_y \rangle$  at time  $t$ .

[4 marks]

c) Find the other eigenvalues of  $\hat{L}_y$  and their eigenvectors in the  $L_z$  basis.

[11 marks]

d) Find the probability of measuring  $L_y = +\hbar$  at time  $t$ .

[4 marks]

**END OF EXAMINATION PAPER**

## PHYSICAL CONSTANTS AND CONVERSION FACTORS

SYMBOL	DESCRIPTION	NUMERICAL VALUE
$c$	Velocity of light in vacuum	$299\,792\,458\text{ m s}^{-1}$ , exactly
$\mu_0$	Permeability of vacuum	$4\pi \times 10^{-7}\text{ N A}^{-2}$ , exactly
$\epsilon_0$	Permittivity of vacuum where $c = \frac{1}{\sqrt{\epsilon_0\mu_0}}$	$8.854 \times 10^{-12}\text{ C}^2\text{ N}^{-1}\text{ m}^{-2}$
$h$	Planck constant	$6.626 \times 10^{-34}\text{ J s}$
$\hbar$	$h/2\pi$	$1.055 \times 10^{-34}\text{ J s}$
$G$	Gravitational constant	$6.674 \times 10^{-11}\text{ m}^3\text{ kg}^{-1}\text{ s}^{-2}$
$e$	Elementary charge	$1.602 \times 10^{-19}\text{ C}$
eV	Electronvolt	$1.602 \times 10^{-19}\text{ J}$
$\alpha$	Fine-structure constant, $\frac{e^2}{4\pi\epsilon_0\hbar c}$	$\frac{1}{137.0}$
$m_e$	Electron mass	$9.109 \times 10^{-31}\text{ kg}$
$m_e c^2$	Electron rest-mass energy	0.511 MeV
$\mu_B$	Bohr magneton, $\frac{e\hbar}{2m_e}$	$9.274 \times 10^{-24}\text{ J T}^{-1}$
$R_\infty$	Rydberg energy $\frac{\alpha^2 m_e c^2}{2}$	13.61 eV
$a_0$	Bohr radius $\frac{1}{\alpha} \frac{\hbar}{m_e c}$	$0.5292 \times 10^{-10}\text{ m}$
Å	Angstrom	$10^{-10}\text{ m}$
$m_p$	Proton mass	$1.673 \times 10^{-27}\text{ kg}$
$m_p c^2$	Proton rest-mass energy	938.272 MeV
$m_n c^2$	Neutron rest-mass energy	939.565 MeV
$\mu_N$	Nuclear magneton, $\frac{e\hbar}{2m_p}$	$5.051 \times 10^{-27}\text{ J T}^{-1}$
fm	Femtometre or fermi	$10^{-15}\text{ m}$
b	Barn	$10^{-28}\text{ m}^2$
u	Atomic mass unit, $\frac{1}{12} m(^{12}\text{C atom})$	$1.661 \times 10^{-27}\text{ kg}$
$N_A$	Avogadro constant, atoms in gram mol	$6.022 \times 10^{23}\text{ mol}^{-1}$
$T_t$	Triple-point temperature	273.16 K, exactly
$k$	Boltzmann constant	$1.381 \times 10^{-23}\text{ J K}^{-1}$
$R$	Molar gas constant, $N_A k$	$8.314\text{ J mol}^{-1}\text{ K}^{-1}$
$\sigma$	Stefan-Boltzmann constant, $\frac{\pi^2}{60} \frac{k^4}{\hbar^3 c^2}$	$5.670 \times 10^{-8}\text{ W m}^{-2}\text{ K}^{-4}$
$M_E$	Mass of Earth	$5.97 \times 10^{24}\text{ kg}$
$R_E$	Mean radius of Earth	$6.4 \times 10^6\text{ m}$
$g$	Standard acceleration of gravity	$9.80665\text{ m s}^{-2}$ , exactly
atm	Standard atmosphere	101 325 Pa, exactly
$M_\odot$	Solar mass	$1.989 \times 10^{30}\text{ kg}$
$R_\odot$	Solar radius	$6.96 \times 10^8\text{ m}$
$L_\odot$	Solar luminosity	$3.84 \times 10^{26}\text{ W}$
$T_\odot$	Solar effective temperature	$5.8 \times 10^3\text{ K}$
AU	Astronomical unit, mean Earth-Sun distance	$1.496 \times 10^{11}\text{ m}$
pc	Parsec	$3.086 \times 10^{16}\text{ m}$
	Year	$3.156 \times 10^7\text{ s}$